





SIMBOL-X is one of the 4 formation flying astrophysics mission studied by CNES in 2005 Gamma WAVE - FOCUSING TELESCOPES IN NUCLEAR ASTROPHYSICS _____ R. Clédassou (CNES); Ph. Ferrando (CEA) _____





Why hard-X rays ?

Non thermal emissions in X and γ rays are unique signatures needed to answer fundamental questions in modern astrophysics :

- How works the dynamics of the universe at all scales ? From star formation to cosmological large structure formation, this is driven by accretion power, particularly on Black Holes, and violent non thermal phenomena (as jets)
- How good are our physics laws in extreme conditions, of gravity, pressure, magnetic field ? Do we need new physics ?
- How and where are accelerated the cosmic rays at the highest energies ?



Sept. 2005

SIMBOL-X Mission



Integral > 15 keV

Why Simbol-X ?

Large FOV instruments have unveiled the richness of the domain, but insufficient for doing the physics, and for seeing obscured sources...



30 degrees



Focusing (telescope !) revolution needed in hard X-rays, like was done in optical and soft X-rays, to get "astronomical" sensitivity and angular resolution, and have access to relevant dynamical scales...

Now feasible by extension of the soft X-ray optics to the hard band, thanks to the long focal length offered by the Formation Flight !

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OBSERVATION AREA AND TECHNICAL CONSTRAINTS



Pointing type XMM and INTEGRAL:

allows a sun pointing:

- Fixed Solar Arrays
- Simplified Thermal Control

the celestial vault is scanned in 4,5 months (\approx 35% at any moment)





Minimal life duration : 2,5 years with 2 years for science

Scientific Orbit : 44 000 km / 253 000 km / 7 sidereal days / Low initial inclination

Has been chosen to give 90% time above 73 000 km which maximizes the science

Satellites ΔV around 500 m/s (hydrazine)

Daily visibility : ~12 hours per station with a maximum of 2 hours hole

At perigee the visibility is permanent (24 hours) for the chosen station









PERTURBATIONS ON HEO ORBIT

Orbit: 46000 – 253000 km - formation inclined 45° / orbit plane









Mirror

Communications strategy

TT&C

Stations ESOC : REDU, KOUROU,

MALINDI, PERTH, MADRID

Detector

- 2 communications modes
 - Housekeeping mode (TC + HKTM + ranging)
 - TMCU download (TC + TM high rate, detector only)
- Access the mirror through ISL in nominal phase
- Direct access to mirror (parallel communications) for:
 - Cruise
 - Commissioning
 - Emergency

TT&C for emergency, cruise and commissioning (Rx always ON)

ISL

ESOC (\emptyset 15 m) S band or X band (ev.35m)

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OPERATIONS : SEVERAL HARD SOURCES (22 ks ; > 1 mCrab)











Formation Flying Requirements

Relative positioning :

- lateral / L.O.S. :+/- 1 cm
- longitudinal : +/- 10 cm
- lateral position knowledge : 0.5 mm (or 3" LOS)
- Constraint on inertial and relative metrology lateral sensor + star tracker with 1 arc sec accuracy range

Mirror attitude control:

- Pointing : 10 arc sec
- Stability : no constraint

Detector attitude control:

- Pointing : 1 degree
- Stability : no constraint
- → Low constraints on attitude control









GNC HARDWARE

	Mirror Spacecraft	Detector spacecraft
Metrology	 standard SST (x2) + 1 gyro bloc RF terminal + 1 antenna corner cubes or diodes 	 Fine SST (1") + standard SST or standard SST + precise gyrometers RF terminal RF + 3 antennas Lateral sensor
Actuation	 Reaction wheels (x4) > 1 Nm.s 5 N Hydrazine thrusters (orbit transfer & maintenance; wheels off-loading) 	 Cold gas thrusters (8 x 2) for attitude & fine position control (5.4 kg) 1 N Hydrazine thrusters (orbit transfer & maintenance; formation slew)

FF equipment : no critical technology in metrology (currently in development or R&D studies) and propulsion (already available)







DETECTOR REFERENCE PLANS



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Synthesis & recommendations

- SIMBOL-X is a "permanent observatory" which propose a great number of observations (~500 per year) and a great variety of these sources.
- The short duration of the observation times (less than a day in general) authorizes a great scheduling flexibility or a short term redefinition of the objectives,
- The <u>substantial</u> spacecraft mass margins allows to envisage several mirror options (with adapted focal between 25 m and 30 m) :
 - Mirror thickness similar to XMM mirror,
 - Thinner mirrors,
 - mirrors with multilayer,

⇒ These options shall have to be studied during phase A and the choice confirmed at the end of this one





Synthesis & recommendations (Cont'd)

- The decrease of the transmission at low energy (20% at 1,5keV) should permit to decrease the necessary resources for the mirror module thermal control (addition of insulating layers at the entrance & exit), it is an important simplification with respect to XMM constraints,
- The necessity of including a collimator in front of the detector (perturbations due to diffuse X-ray emission), led to deep modifications of the detector satellite configuration and at a lower level of the mirror satellite one. Nevertheless, a broad space for optimization of the satellite configurations exists,
- In addition, a fairly detailed definition of the detector payload and a strong technical support by CEA allowed to identify suitably the constraints imposed by this one on the space segment.

⇒ The result of this is that the mission and space system are, then, strong & robust (detection function).