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Positrons in Astrophysics – March 2012







Basic processes of positron and positronium physics

Measurements, results and insights; from the bulk to beams and traps

91% H (or H₂); 9% He: 0.1% "metals"







Partial Chronology - Experiment



- 1949:- Discovery of Ps
- 1949-50:- Discovery of "ACAR"
- 1955:- Discovery of antiproton
- 1950's:- First attempts to make positron beams
- 1963:- Anomalous annihilation rates on some molecules
- 1968:- Ps formation in powders
- 1971-2:- First lab-based beam
- 1978-88:- Understanding of beam production
- 1981:- Discovery of Ps-
- 1986:- Solid neon moderator
- 1988:- First positron plasma in lab. (accumulation device)
- 1995-6:- Discovery of antihydrogen (relativistic)
- 2000's:- Development of robust plasma manipulation techniques for positrons
- 2002:- First cold antihydrogen
- 2007:- Discovery of Ps_2
- 2010:- Trapped antihydrogen
- 2011-12: First measurement of antihydrogen transition









Direct Annihilation: $e^+ + A \longrightarrow 2\gamma + A^+$

Cross section at low energies = $\pi r_0^2 c Z_{eff} / v$

 $Z_{eff} \approx 9$ for H and 4 for He. Can be considerably larger for other species (e.g. classes of organic molecules)









Positronium Formation:

 $e^+ + A \longrightarrow Ps + A^+$

Threshold at positron kinetic energy $E_{ps} = (E_i - 6.8/n^2) \text{ eV}$ for formation into a Ps state with principal quantum number, n

Cross section typically rises rapidly from threshold, peaks within around 10-15 eV and falls off at higher energies. It is a small fraction of the total inelastic scattering cross section above 100 eV.









Positronium Formation:

 $e^+ + A \longrightarrow Ps + A^+$

Some examples of E_{ps}

H (~ 6.8 eV); H₂ (~8.6 eV); He (17.8 eV)

However, there are many species for which $E_{ps} < 0$ (Li, Na, K, Rb, Cs, Ca, Sc, Ti, V, Cr, Ga, Ba, La, Ce, Pr, Nd, Pm, Sm, Eu, Gd, Tb, Dy, Ho, Er, Tm, Yb, Lu, Hf, Fr, Ra, Ac and several more where E_{ps} is within 0.5 eV of zero (Mn, Mo, Tc, Ru, Rh, Ag, Sn, Pb, Bi)









Positronium Formation:

 $e^+ + A \longrightarrow Ps + A^+$

For those many species for which $E_{ps} \approx 0$, little is known about the size of the cross section, except that it will be very large and diverge as E tends to 0.

Excited states of Ps may also be preferred in these cases.









Excitation: $e^+(E) + A \longrightarrow e^+(E - E_{ex}) + A^*$

For a threshold at E_{ex}

Cross section typically rises smoothly from threshold; usually only a few excited states make important contributions to the overall inelastic scattering cross section.









Ionization: $e^+(E) + A \longrightarrow e^+(E - \Delta E) + A^+$

Threshold at E_i

Cross section, $\sigma_i(E)$, typically rises smoothly from threshold to become the dominant process when Ps formation ceases to occur.

 ΔE around 20-40 eV. Need to know single differential cross section, $d\sigma_i(E)/d(E - \Delta E)$ to model energy loss due to this process.

Higher order ionization process (for atoms other than H) can be ignored.









Annihilation: $Ps(1^{3}S) \longrightarrow 3\gamma$: Vacuum lifetime around 142 ns

Ps (1¹S) \longrightarrow 2 γ : Vacuum lifetime about 125 ps

Excited S states have lifetimes longer by a factor of n^{3.} Higher angular momentum states are essentially stable against annihilation, but radiatively decay with lifetimes expected to be around twice those of corresponding state in H.









Scattering:
Ps + A
$$\longrightarrow$$
 Ps + A
 $2\gamma + e^- + A^+$
Ps^{*} + A
Ps + A^{*}
 $e^+ + e^- + A$
Ps + $e^- + A^+$

The cross sections for some of these processes are known for a few targets (typically the noble gases). However, with typical collision frequencies in astrophysical environments much lower than the inverse lifetime, these processes don't seem to important.













Time-honoured traditional positron lifetime method.

Measured parameters are Z_{eff} , $_1Z_{eff}$ (pickoff parameter for Ps) and the positronium fraction, F.

Noble gases and a variety of small molecules



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Measurements and Results - Bulk





FIGURE 2 - Possible fates of positrons in a gas [6].



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Measurements and Results - Bulk





"Shoulder" at short times due to annihilation before thermalization. Effect mostly below about a couple of eV.

Thermalization time about 350 nsamagat; approx. 1.3 x 10^{16} sm⁻³

All gas atoms with an appreciable E_{ps} will have similar thermalization times.



Positrons in Astrophysics – March 2012 <u>Positron and Positronium Scattering and Annihilation in Atomic and Molecular Systems</u>



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Thermalization times of positrons in molecular gases

Ilham Al-Qaradawi[†], Michael Charlton[‡], Ivan Borozan and Ralph Whitehead§

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Various amounts of N_2O added to Ar. Max. amount about 1%.





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Measurements and Results - Bulk



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Thermalization times of positrons in molecular gases

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Received 27 March 2000, in final form 31 May 2000

1ns-amagat $\approx 2.7 \text{ x } 10^{10} \text{ scm}^{-3}$ \dots or about 800 y at 1 cm⁻³

Some examples: (in ns-amagat)

H_2	2.4
N_2	14
CO	1.1
CO_2	0.1
SF ₆	0.04
He	1700
Ar	360

However – not strictly applicable to astrophysical context ...



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Positronium Fraction, F

Global parameter – fraction of positrons forming Ps after slowing down in the gas.

Not equivalent to astrophysical measurements as Ps collisions lead to break-up when Ps kinetic energy exceeds 6.8 eV. Apply so-called Ore model (uniform E).

$$F_{min} = (E_{ex} - E_{ps})/E_i$$
 $F_{max} = 6.8/E_i$









Positronium Fraction, F Some examples

Species	F_{min}	F _{max}	F
Н	0.25	0.50	?
Не	0.14	0.28	0.23
Ne	0.09	0.32	0.26
Ar	0.17	0.43	0.33
Kr	0.20	0.49	0.11
Xe	0.26	0.56	0.03
H_2	0.18	0.44	0.32
N_2	0	0.44	0.19
$-CH_4$	0.18	0.52	0.40



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Positronium Fraction, F How about gas mixtures?

Imagine adding another atomic species with a very low $E_{\rm Ps}$ to a H sample \ldots

The H Ore gas is from 6.8-10.2 eV. Any positron which falls to a very low energy will have a good chance of forming Ps with the added gas, as the only competition is energy loss via momentum transfer in e⁺-H collisions.

Such a method was used to deduce Ps formation cross sections (with lots of assumptions) before beam measurements became available.









Reasonably mono-energetic beams of positrons available since ~ 1970. ($\delta E \sim 1 \text{ eV}$) Cross sections for a number of processes and targets exist.

What can trap-based devices deliver?

Trapped clouds and plasmas for experimentation (annihilation, antihydrogen)

Ability to manipulate in position and time to produce controlled plasmas and/or large instantaneous fluxes of positrons (and hence positronium) – on target, or as a target (molecular Ps, Ps spectroscopy, positron lifetimes in solids)

Slow-release narrow energy-width beams for scattering and annihilation studies with unprecedented energy resolution









H and He are best known from theory. H more-or-less exact at all important energies for the major processes.

 $\sigma_{\rm T}$:: total scattering cross section

 σ_{Ps} :: Ps formation

 σ_{ex} :: excitation

 σ_i :: ionization

Also "total ionization" :: $(\sigma_{Ps} + \sigma_i)$











Comparing electron and positron total scattering cross sections for He

The underlying interactions:

	Positron
Static	Repulsive
Polarization	Attractive
Exchange	No

Electron Attractive Attractive Yes



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Positrons in Astrophysics – March 2012 Positron and Positronium Scattering and Annihilation in Atomic and Molecular Systems



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PHYSICAL REVIEW A 80, 032710 (2009)

High-resolution positron scattering from helium: Grand total and positronium-formation cross sections

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FIG. 3. (Color online) (a): Present grand total cross section (\bullet) and the experimental results of Brenton *et al.* [6] (\blacktriangle), Kauppila *et al.* [9] (\blacklozenge), Canter *et al.* [7] (\blacksquare), Stein *et al.* [10] (\diamondsuit), and Coleman *et al.* [8] (\square). (b): Present grand total cross section (\bullet) compared with theoretical calculations of Baluja *et al.* [18] (--), Cheng *et al.* [19](--) Campbell *et al.* [20] ($-\cdots$) and present CCC from $l_{max}=8$ calculation (-) (see text).



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PHYSICAL REVIEW A 80, 032710 (2009)



FIG. 4. (Color online) (a): Comparison between present positronium-formation cross section results (\bullet), and those of Murtagh *et al.* [23] (\blacktriangle), Overton *et al.* [21] (\bigstar), Fromme *et al.* [24] (\bigcirc), Diana *et al.* [25] (\square), and Fornari *et al.* [22] (\blacksquare). (b): Present positronium-formation cross section results (\bullet) compared with theoretical calculations of Cheng *et al.* [19](--), Campbell *et al.* [20] (----), McAlinden *et al.* [27] (---), and present CCC theory (-).



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INSTITUTE OF PHYSICS PUBLISHING JOURNAL OF PHYSICS B: ATOMIC, MOLECULAR AND OPTICAL PHYSICS

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Positron-impact ionization and positronium formation from helium

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INSTITUTE OF PHYSICS PUBLISHING

JOURNAL OF PHYSICS B: ATOMIC, MOLECULAR AND OPTICAL PHYSICS

J. Phys. B: At. Mol. Opt. Phys. 35 (2002) 2525-2540

PII: S0953-4075(02)32270-3

Total positron-impact ionization and positronium formation from the noble gases

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Figure 2. Total and direct ionization cross-sections for positron impact on Ar: \diamond —present total ionization cross-section; \Box —direct ionization cross-section of Moxom *et al* (1996); \bigtriangledown —total single-ionization cross-section of Bluhme *et al* (1999b); \bigcirc —double-ionization cross-section of Bluhme *et al* (1999b); \bigcirc —double-ionization cross-section. The ionization cross-section for electron impact (Sorokin *et al* (2000); solid curve) used for normalization purposes is also shown.



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Fig. 4.20. Positronium formation in positron-hydrogen scattering. Experiment: G, Zhou et al. (1997); o, Hofmann et al. (1997); o, Kara (1999). Theory: -----, Brown and Humberston (1985); ----, Kernoghan et al. (1996); ----, Mitroy (1996); — · —, Igarishi and Toshima (1994); · · · · ·, Higgins and Burke (1993); $-\cdots$, Janev and Solov'ev (1998).



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collisions

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Ionization in e+-H collisions Letter to the Editor

L486



Figure 3. Positron impact ionization cross sections for aromic hydrogen: O, present results; \bullet , Spicher *et al* (1990). The line shows the e^- data of Shah *et al* (1987) for comparison.



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PHYSICAL REVIEW A 72, 062713 (2005)

Positron-impact ionization, positronium formation, and electronic excitation cross sections for diatomic molecules

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(Received 1 September 2005; published 21 December 2005)





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FIG. 11. Integral cross sections in N₂ for the excitation of the $a^{1}\Pi$ states by (\bullet) positrons and (\bigcirc , \square) [24,25] electrons. Also



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PRL 102, 133202 (2009) PHYSICA

PHYSICAL REVIEW LETTERS

week ending 3 APRIL 2009

Excited-State Positronium Formation from Helium, Argon, and Xenon

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J. Phys. B: At. Mol. Opt. Phys. 29 (1996) 3971–3987. Printed in the UK Positron scattering by rubidium and caesium

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Key: short dash, n = 1 Ps formation; long dash n = (1 + 2); dot dash n = (1 + 2 + 3); full, total as $n = (1 + 2 + 3) + n^3$ addition for $n \ge 4$. Experimental data from Detroit group; Surdutovitch *et al.* Phys. Rev A 53 (1996) 2861











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Fragmentation of Positronium in Collision with He Atoms

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FIG. 1. Schematic diagram of the Ps beam.



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Break-up cross section in Ps-He scattering versus two theories; line -Blackwood *et al.* Phys Rev A 60 (1999) 4454 and dash – Biswas and Adhikari, Phys Rev A 59 (1999) 363



FIG. 3. Longitudinal energy spreads of the positrons released from Ps breakup.



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Electron-Like Scattering of Positronium

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PHYSICAL REVIEW A 81, 012715 (2010)

Positronium cooling in porous silica measured via Doppler spectroscopy

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243 nm linewidths versus positron implantation energy into a porous sample.

Gaussian fits are shown

Probe laser bandwidth (0.02 nm)



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PRL 108, 043401 (2012) PHYSICAL REVIEW LETTERS

week ending 27 JANUARY 2012

S Efficient Production of Rydberg Positronium

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Laser Pulse

Top View

+3.7kV

Skimmer

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Photodetachment of Positronium Negative Ions

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(b)

Grid B Grid A BUT now its e[†]pulse 4.2keV Na/W DS negative ion ... O +2.7kV B e

GND

week ending 15 APRIL 2011

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Most processes of interest have cross sections that are reasonably well-known, or better.

If you need anything, ask ...

Experimenters and theorists in the positron and positronium field have well-developed sets of tools to hand!

