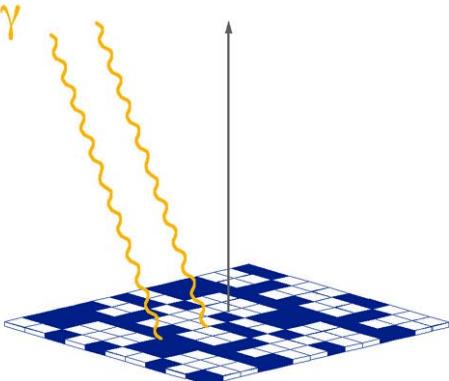
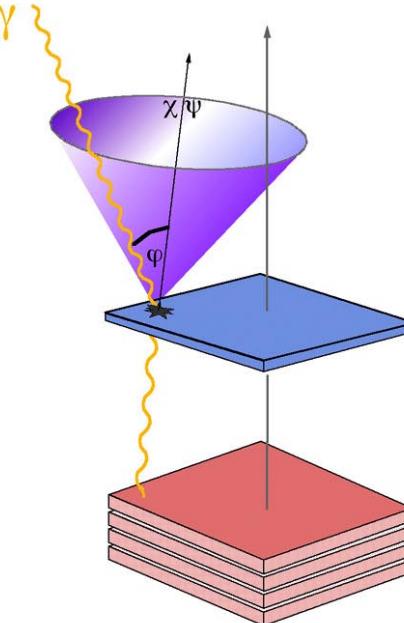
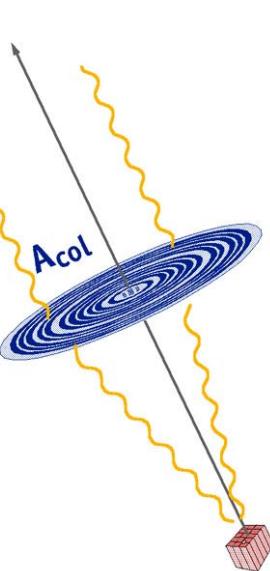
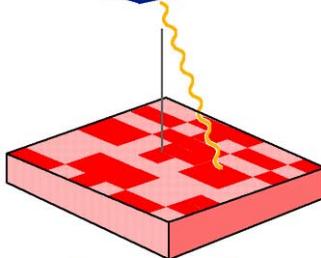
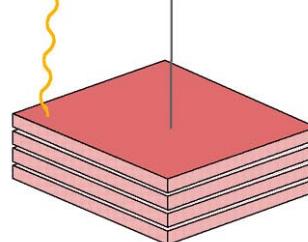


Focaliser des Rayons Gamma ? Pourquoi ?

	modulating aperture systems	Compton telescopes	crystal lens telescopes
aperture / effect	geometric optics absorption	quantum optics incoherent scattering	wave optics coherent scattering
aperture system			
detector	 $A_{\text{det}} = A_{\text{col}}$	 $A_{\text{det}} = A_{\text{col}}$	 A_{det}
signal S ~	A_{col}	A_{col}	A_{col}
background B ~	$V_{\text{det}} \sim A_{\text{det}} = A_{\text{col}}$	$V_{\text{det}} \sim A_{\text{det}} = A_{\text{col}}$	$V_{\text{det}} \sim A_{\text{det}} \ll A_{\text{col}}$
S/B ≈	$\text{const}(A)$	$\text{const}(A)$	$A_{\text{col}}/A_{\text{det}}$

PvB 2001
©

Focaliser des Rayons Gamma ?

... the inability to reflect or deflect individual photons makes the **concentration of a gamma-ray beam impossible.**

A. J. Dean, *Nuclear Instruments and Methods in Physics Research* **221**, 1984

Focusing gamma rays seems out of the question since their wavelengths (less than 0.01 angstrom) are smaller than the distance between atoms in solids.

Giovanni F. Bignami, *Sky & Telescope*, October 1985

Higher-energy X-ray photons can pass through a lens, but since they undergo no significant deflection, **no focusing** can take place.

Gerald K. Skinner, *Scientific American*, August 1988

... gamma-rays can not be focused. They are scattered incoherently and the direction of the scattered electrons are lost.

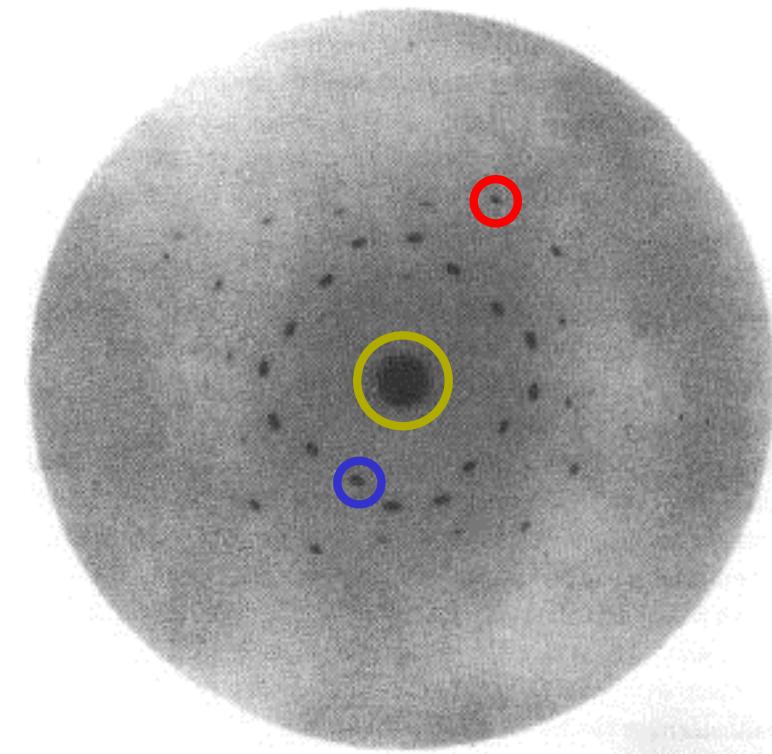
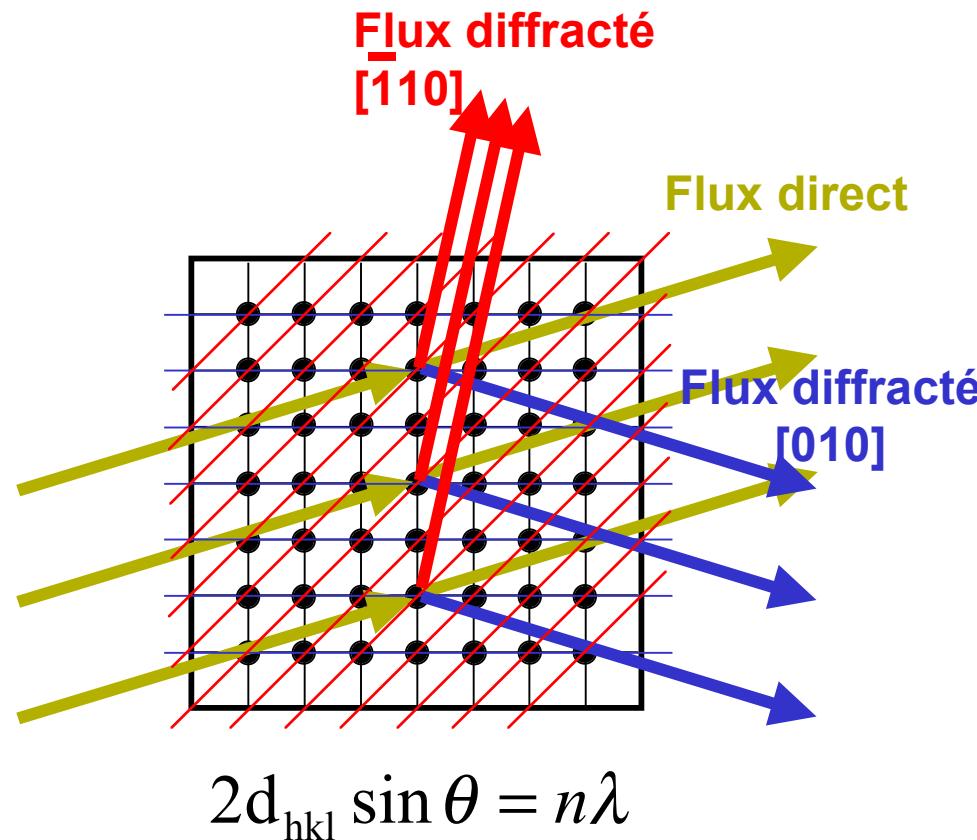
von Ballmoos et al., *Astron. Astrophys.* **221**, 396, 1989

Focaliser des Rayons Gamma !

“Tout le monde croyait que c’était impossible,
sauf un imbecile, qui ne le savait pas et qui l’a fait”

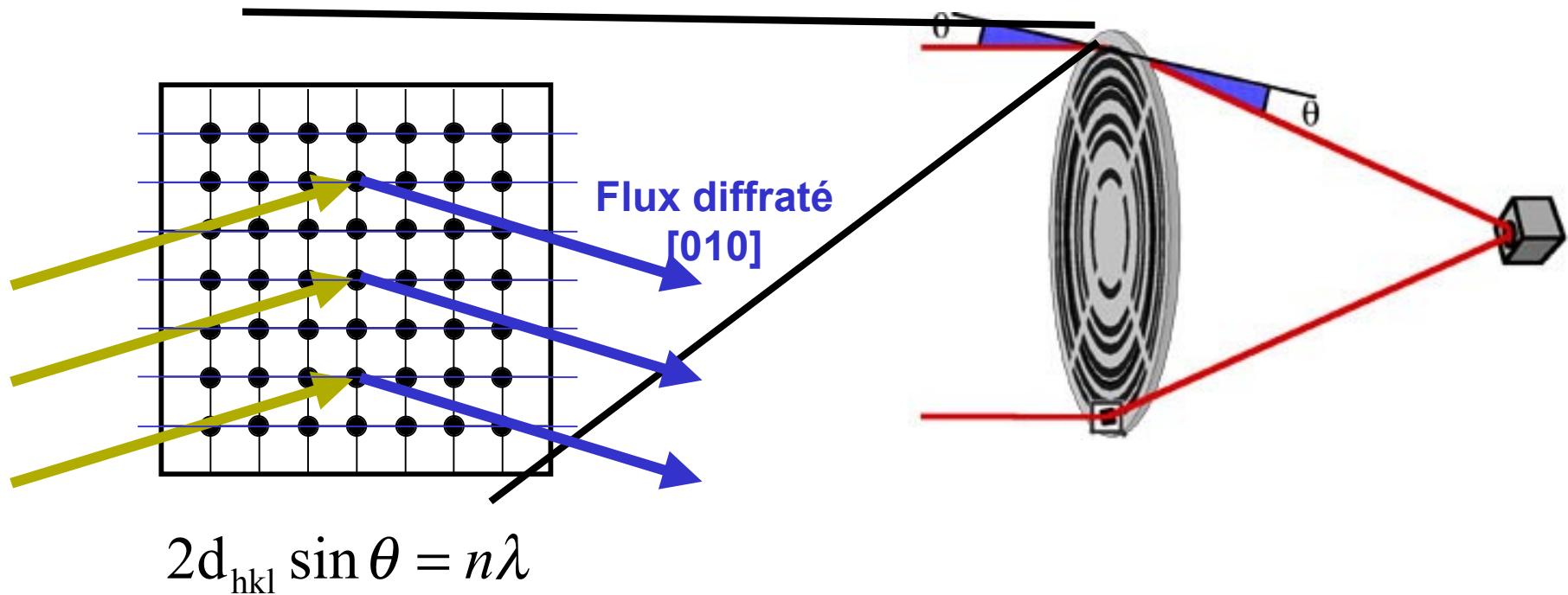
Marcel Pagnol, 1895-1974.

Focaliser les rayons gamma - comment ?

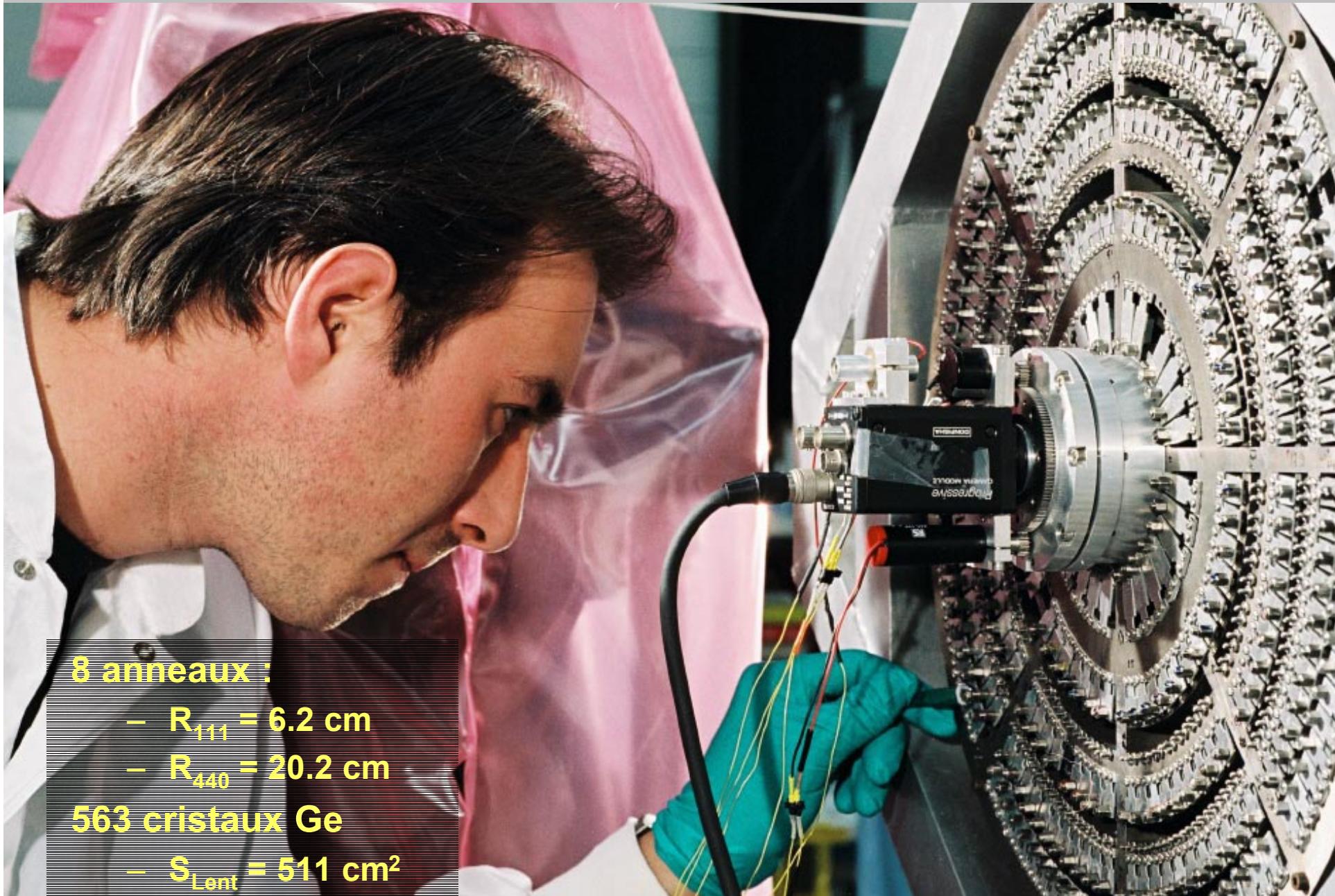


Laue, Friedrich et Knipping, 1912

Focaliser les rayons gamma - comment ?



CLaire - première lumière d'une lentille gamma



8 anneaux :

- $R_{11} = 6.2 \text{ cm}$
- $R_{440} = 20.2 \text{ cm}$

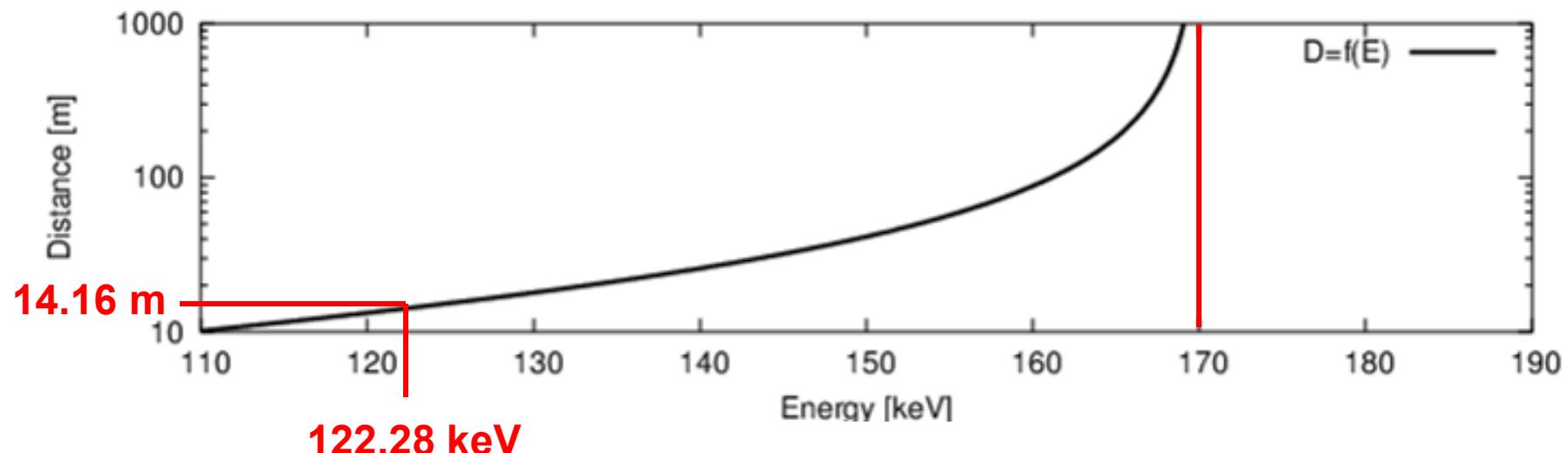
563 cristaux Ge

- $S_{\text{Lent}} = 511 \text{ cm}^2$

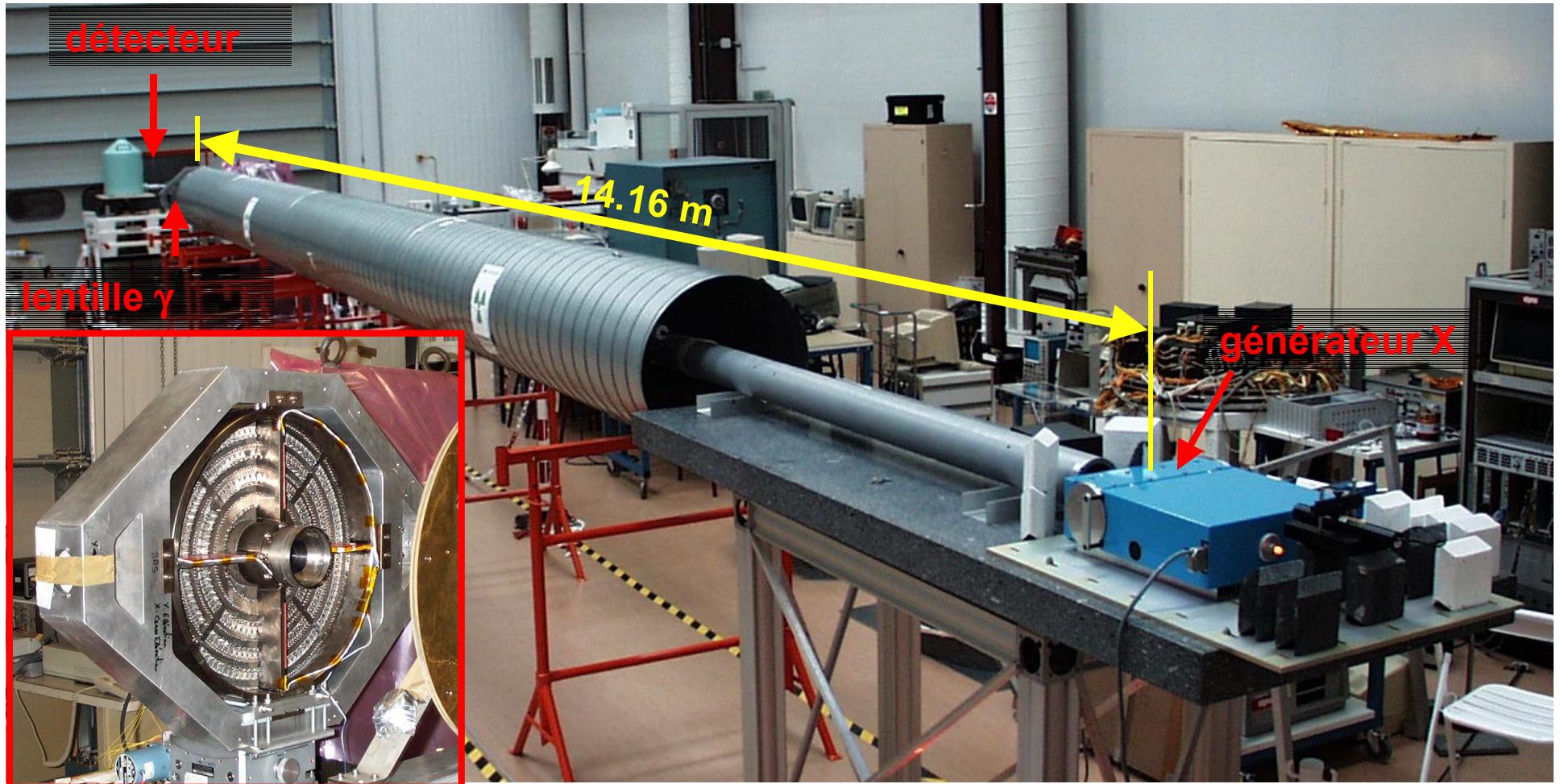
CLaire : principe de réglage de la lentille

Relation entre distance de la source et énergie diffractée :

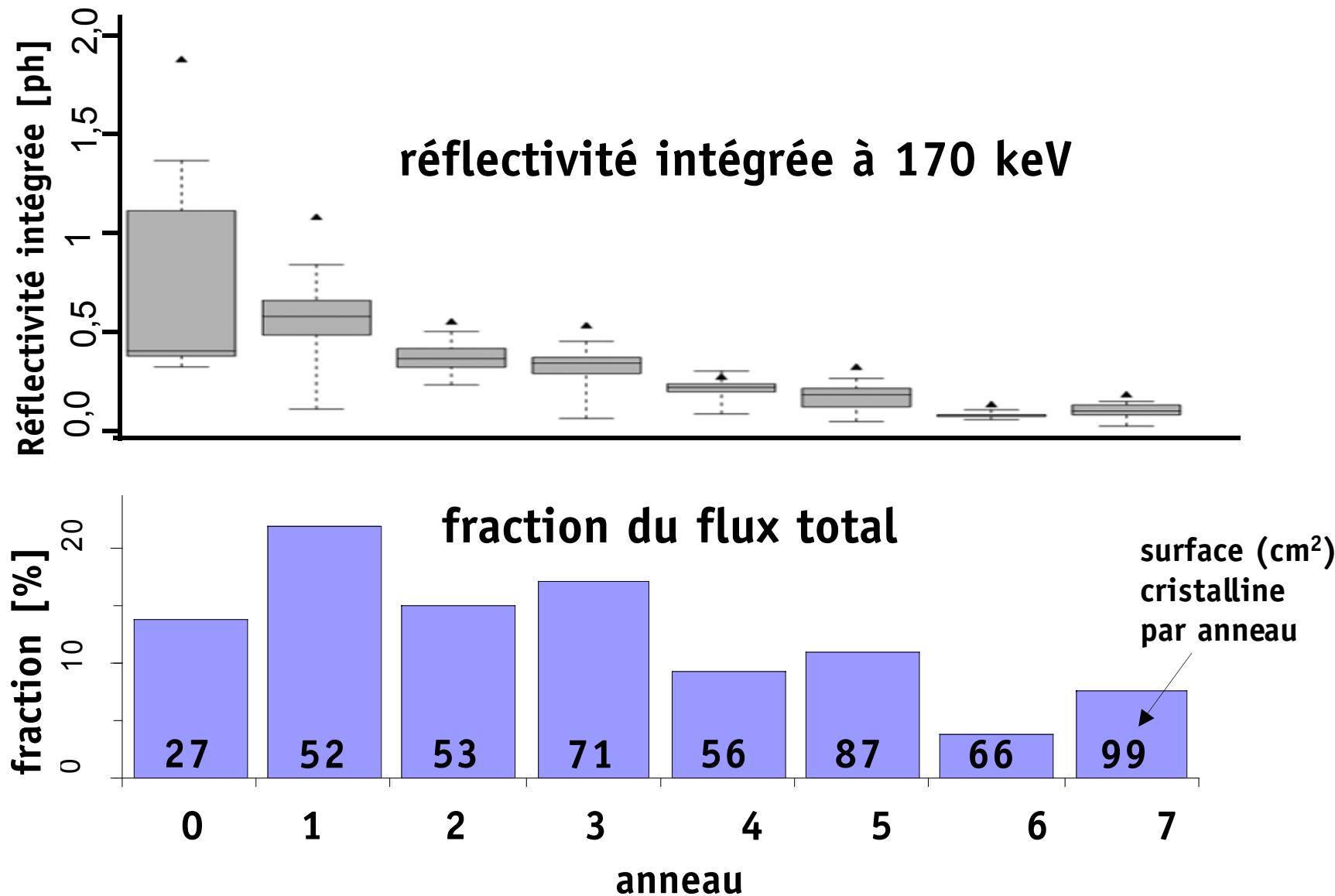
$$\frac{100 \text{ keV}}{E} = \frac{100 \text{ keV}}{E_\infty} + 0.3251 \left(\frac{10 \text{ m}}{D} \right) \quad \text{avec } E_\infty = 170 \text{ keV}$$



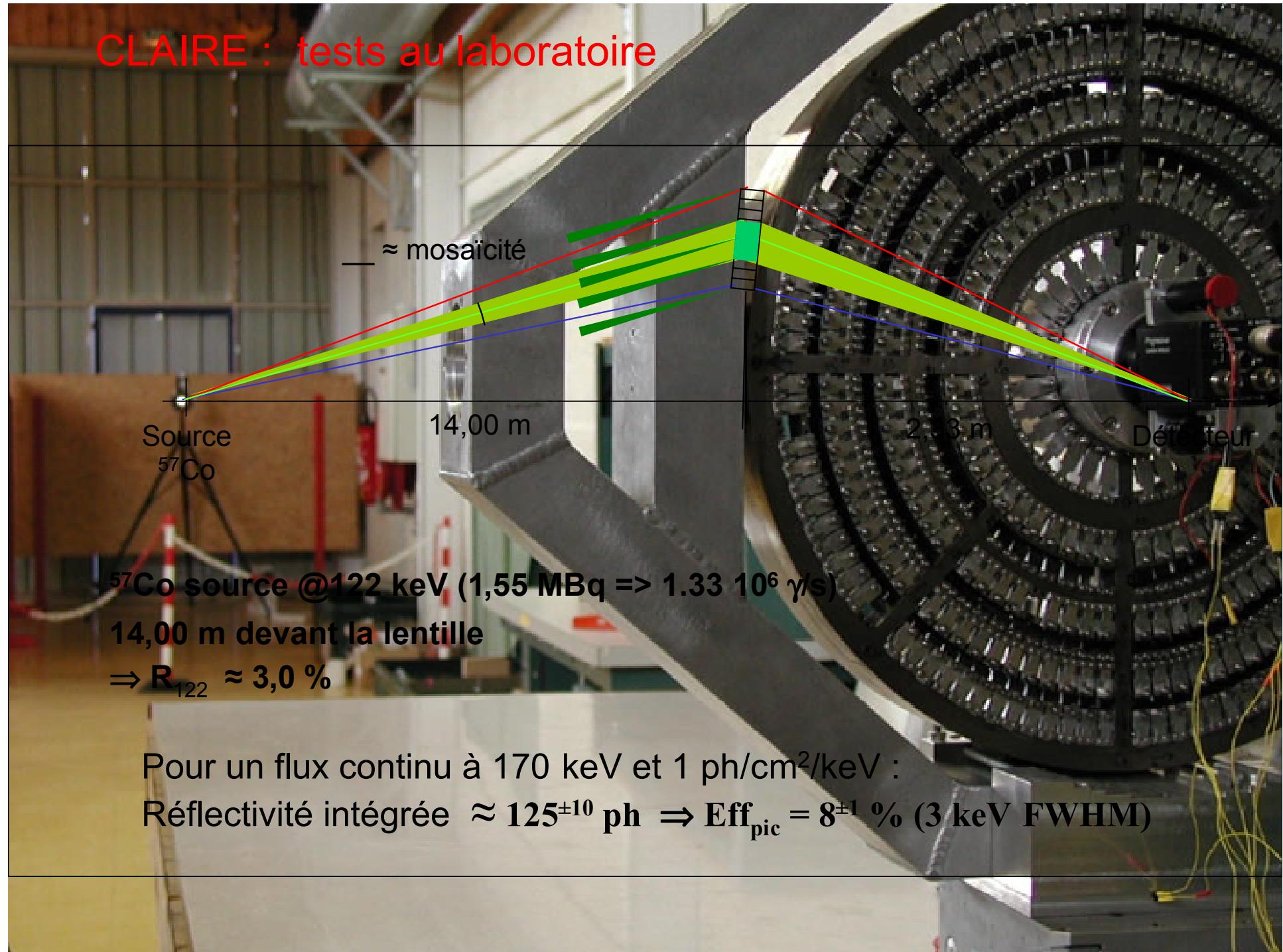
CLAIRE : le banc optique gamma au CESR



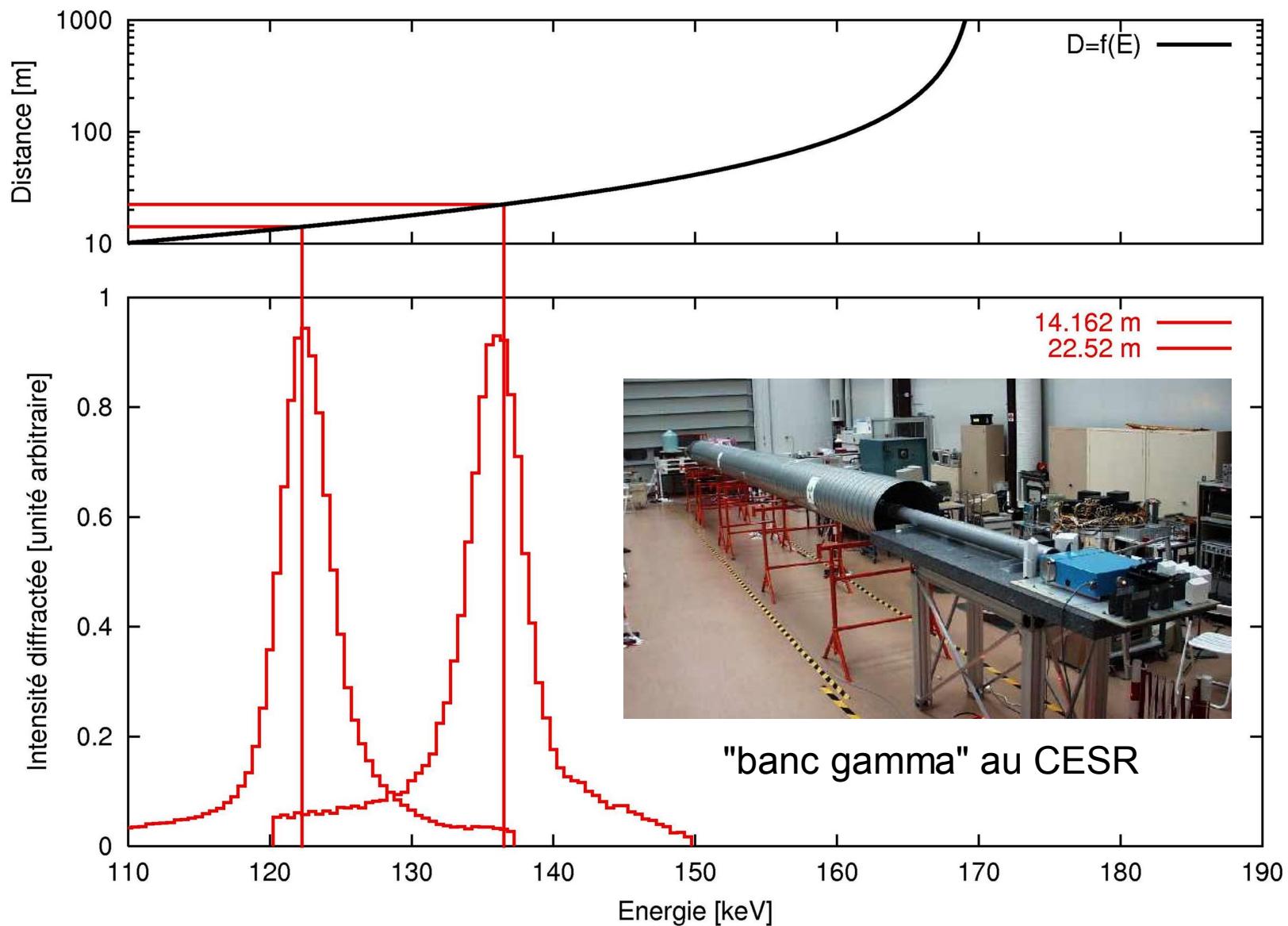
CLaire : qualité des cristaux individuels



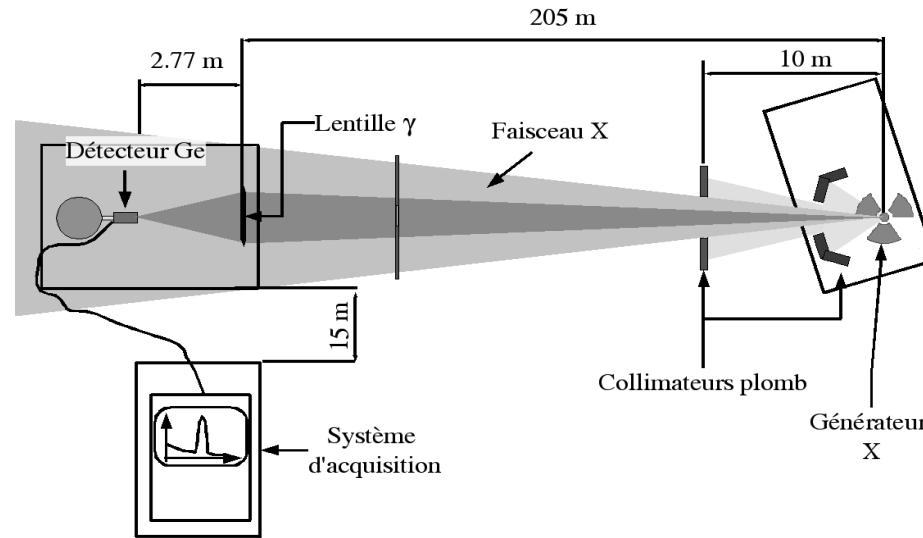
CLAIRES : tests au laboratoire



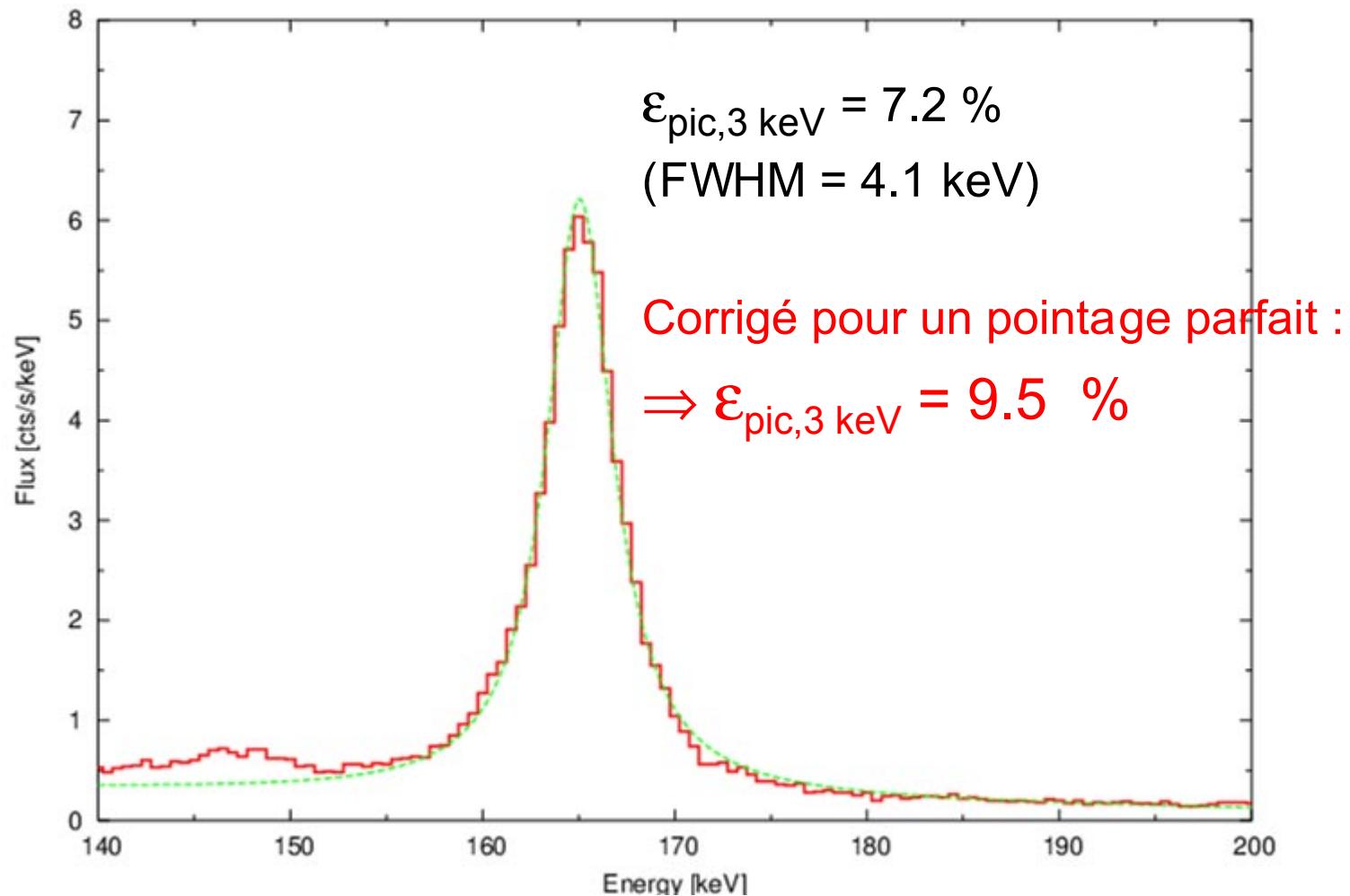
CLAIRE : tests au laboratoire ... et au delà



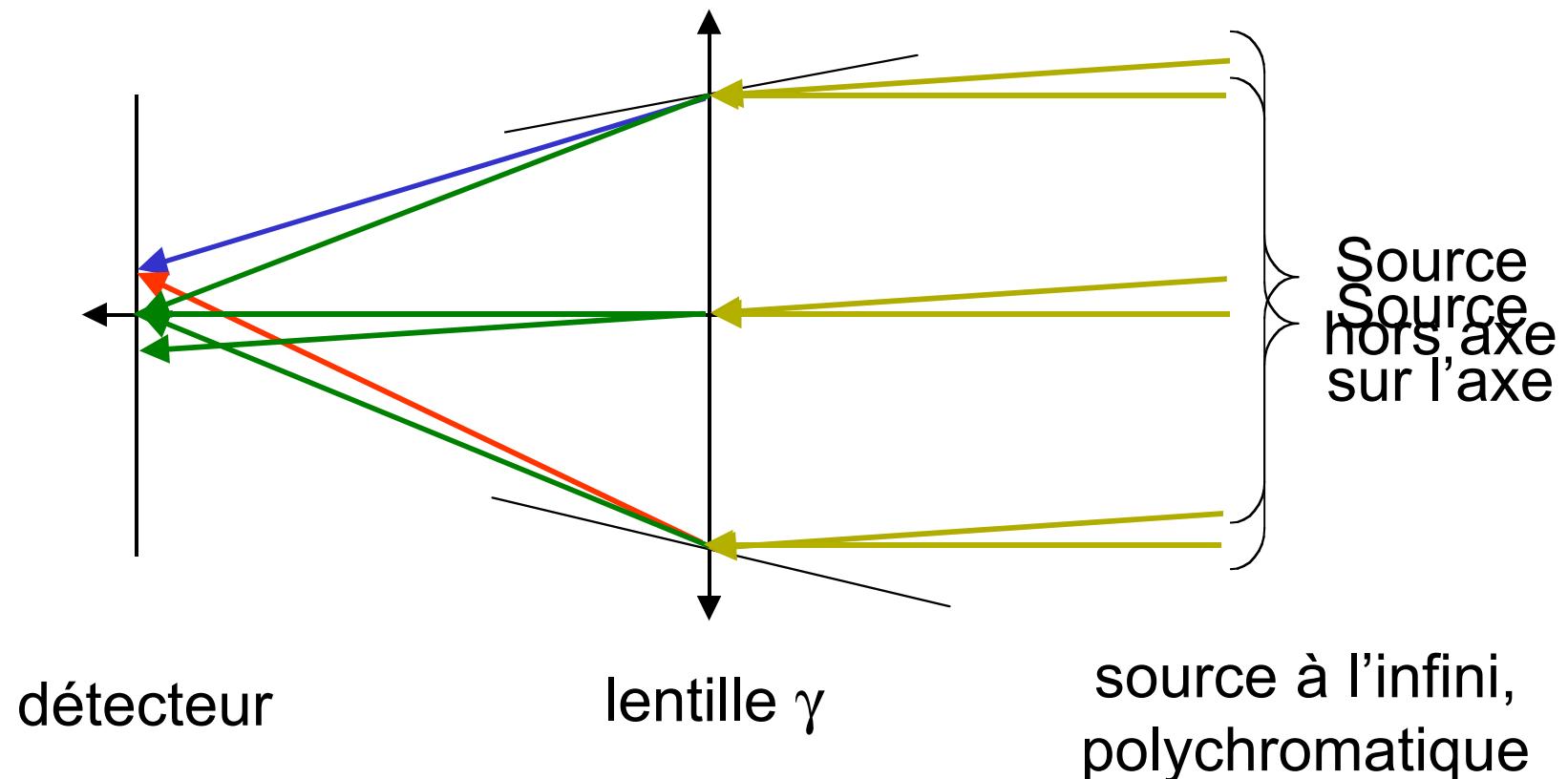
CLAIRE TGD : une source proche de l'"infini" ...



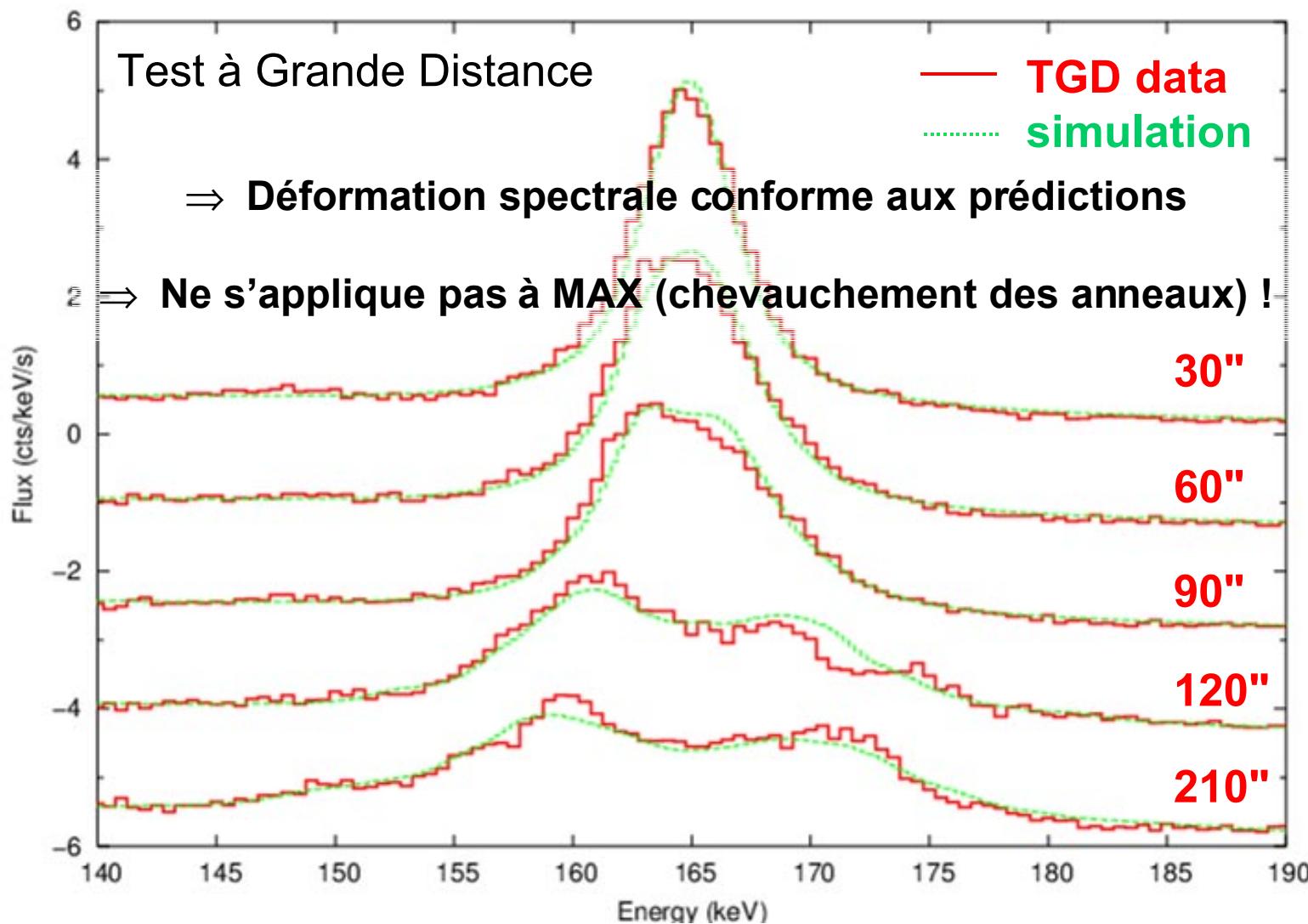
CLAIRES TGD - efficacité de diffraction



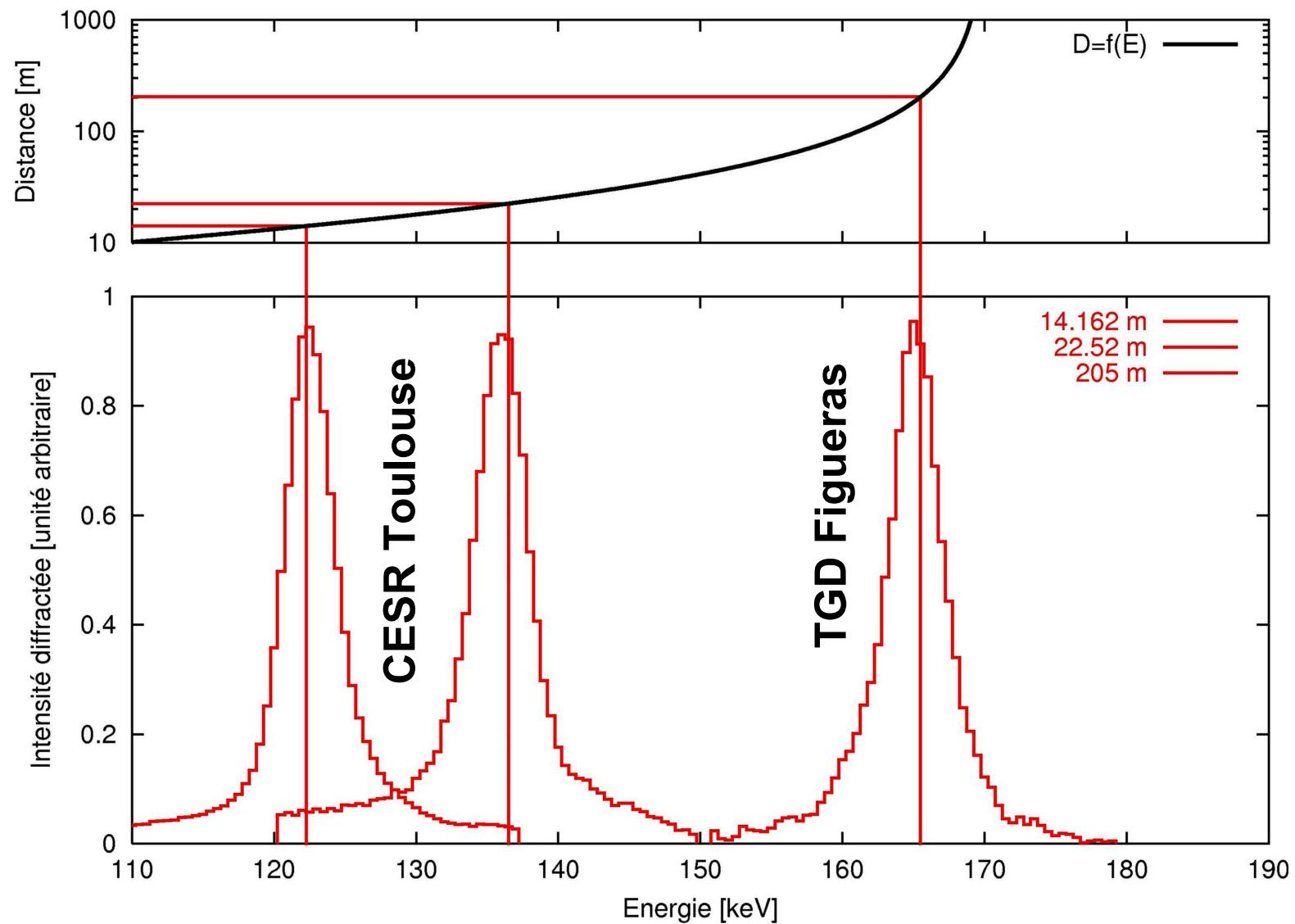
Réponse instrumentale de la lentille - principe



CLAIRE TGD : réponse hors axe



CLAIRE TGD : 14 m, 22.5 m ... et 205 m



CLAIRE 2001



but : démontrer le principe de lentille γ sur un objectif astrophysique

- | | |
|--------------|---|
| Lancement | : 14 juin 2001, 6h15 UT, base CNES de Gap-Tallard |
| Ballon | : Zodiac Z600 (600.000 m ³) |
| Altitude | : > 41 km (3,8 g/cm ² atmosphère résiduelle), pendant 5h 30' |
| Récupération | : 14 juin 2001, 15 h UT, Bergerac, Aquitaine |

CLAIRE 2001 : la nacelle

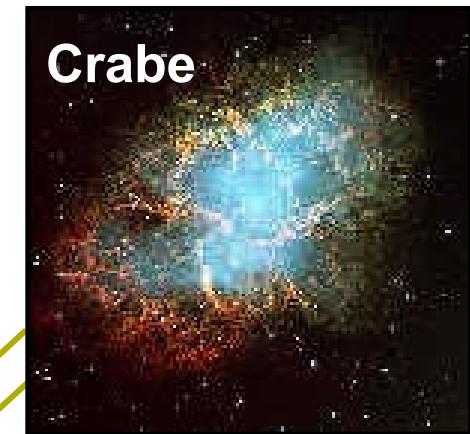
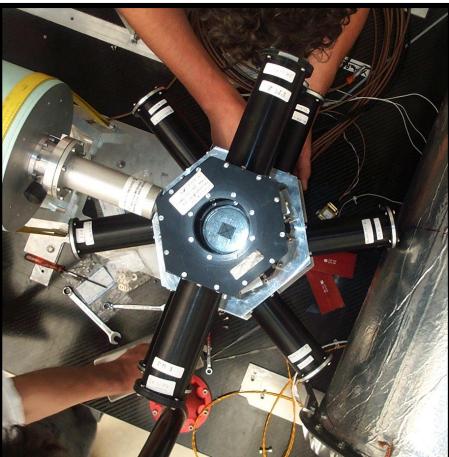
chandelle standard, polychromatique
proche du Soleil ($\sim 1^\circ$ le 15 Juin)

Détecteur

- matrice 3x3
- Ge haute pureté
 $1,5 \times 1,5 \times 4$ cm

Blindage actif

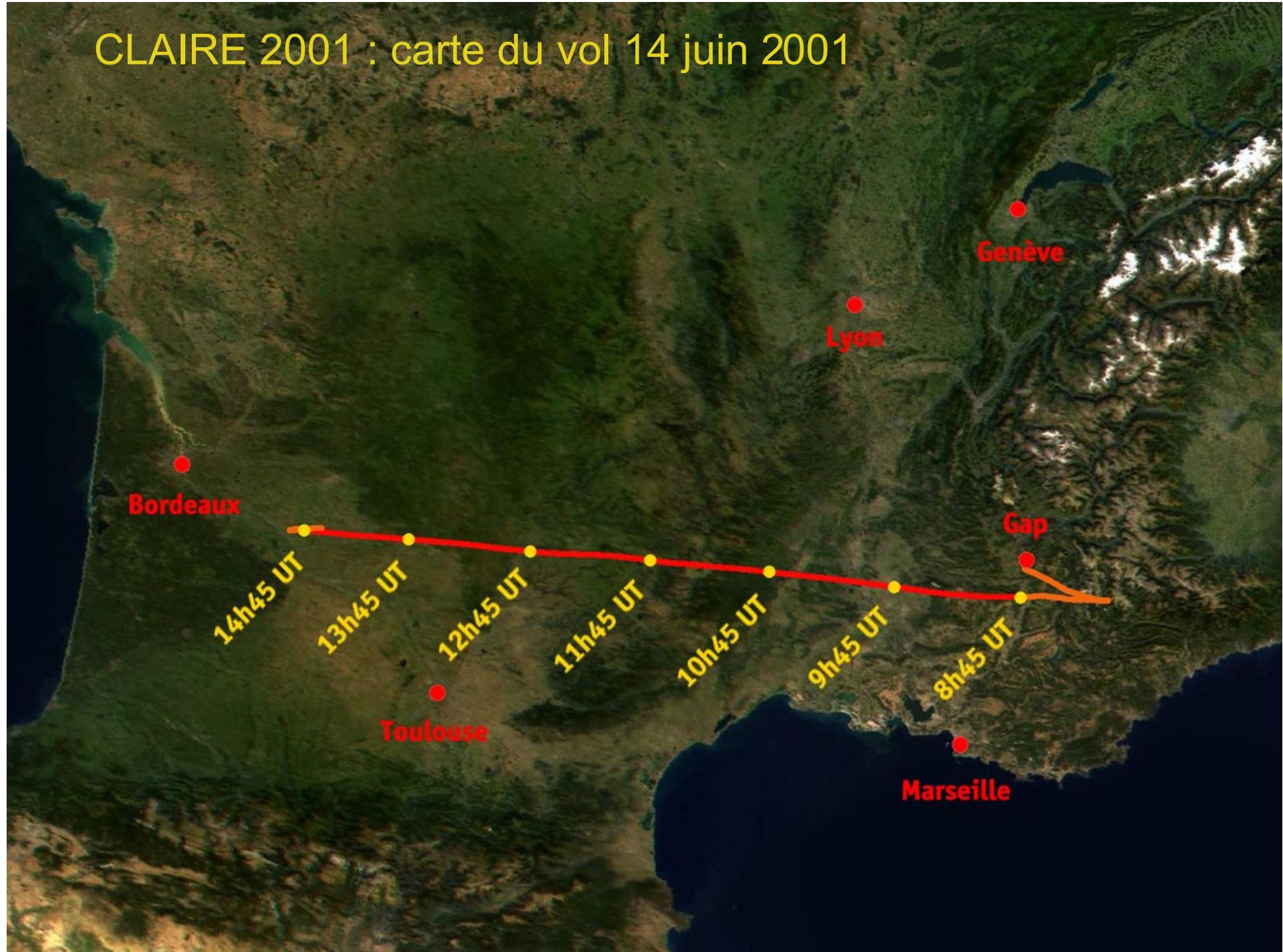
- Cs
- BGO



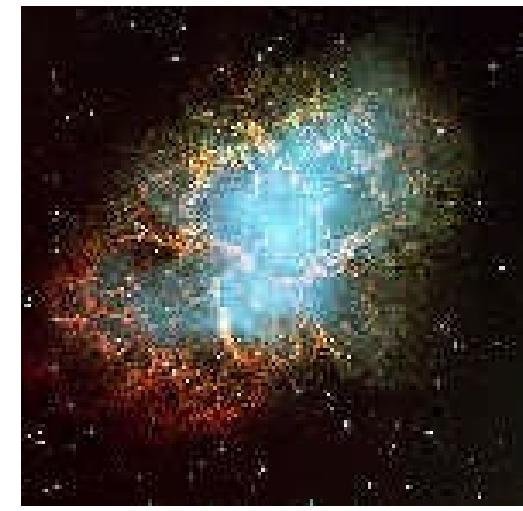
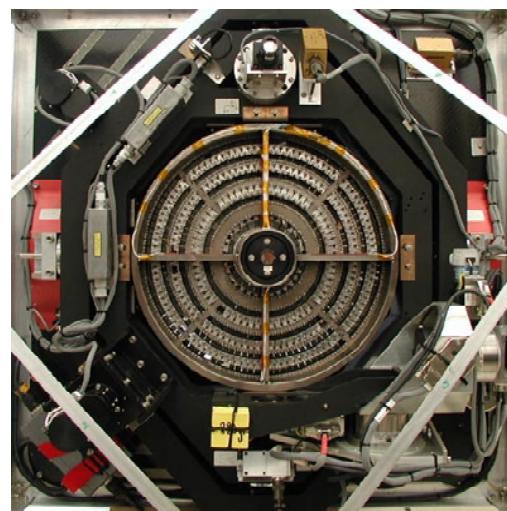
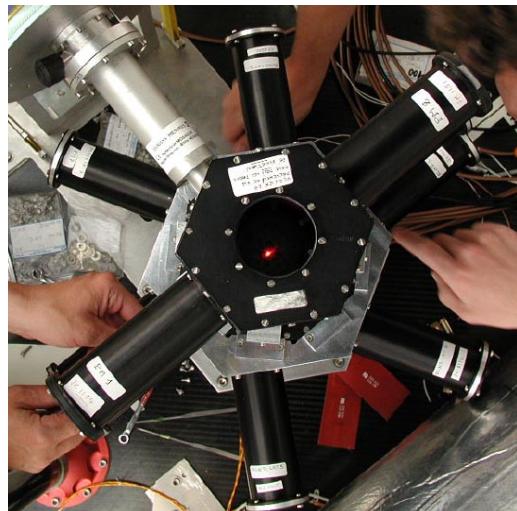
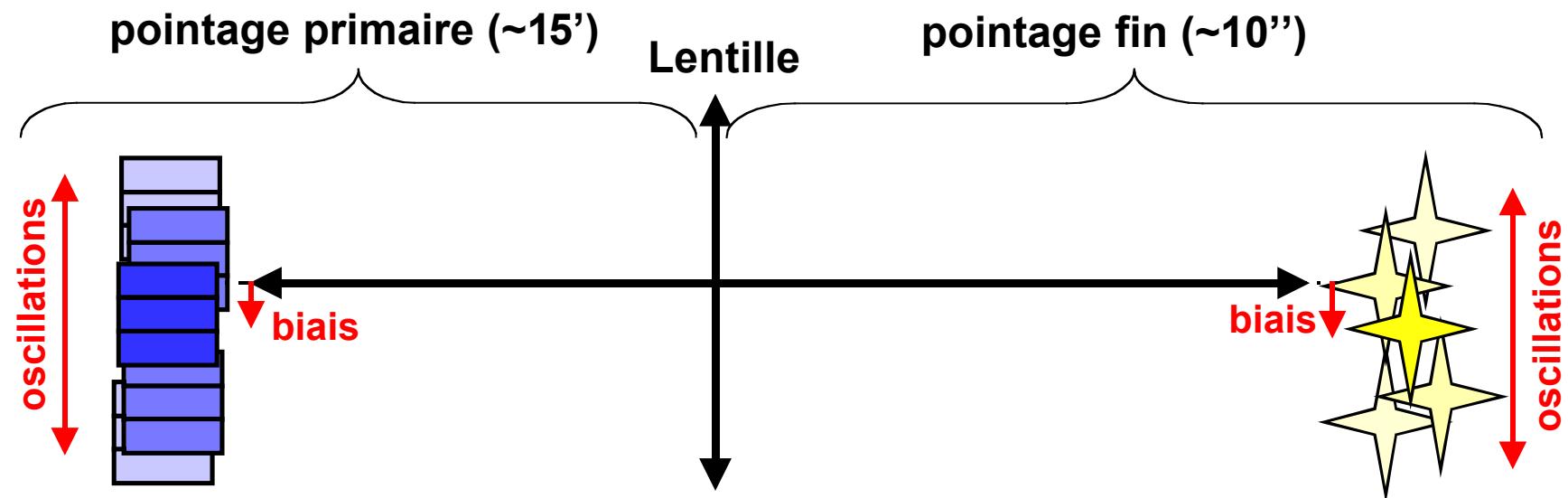
lentille γ

- 563 cristaux
- $E = 170$ keV
- FWHM ~ 3 keV

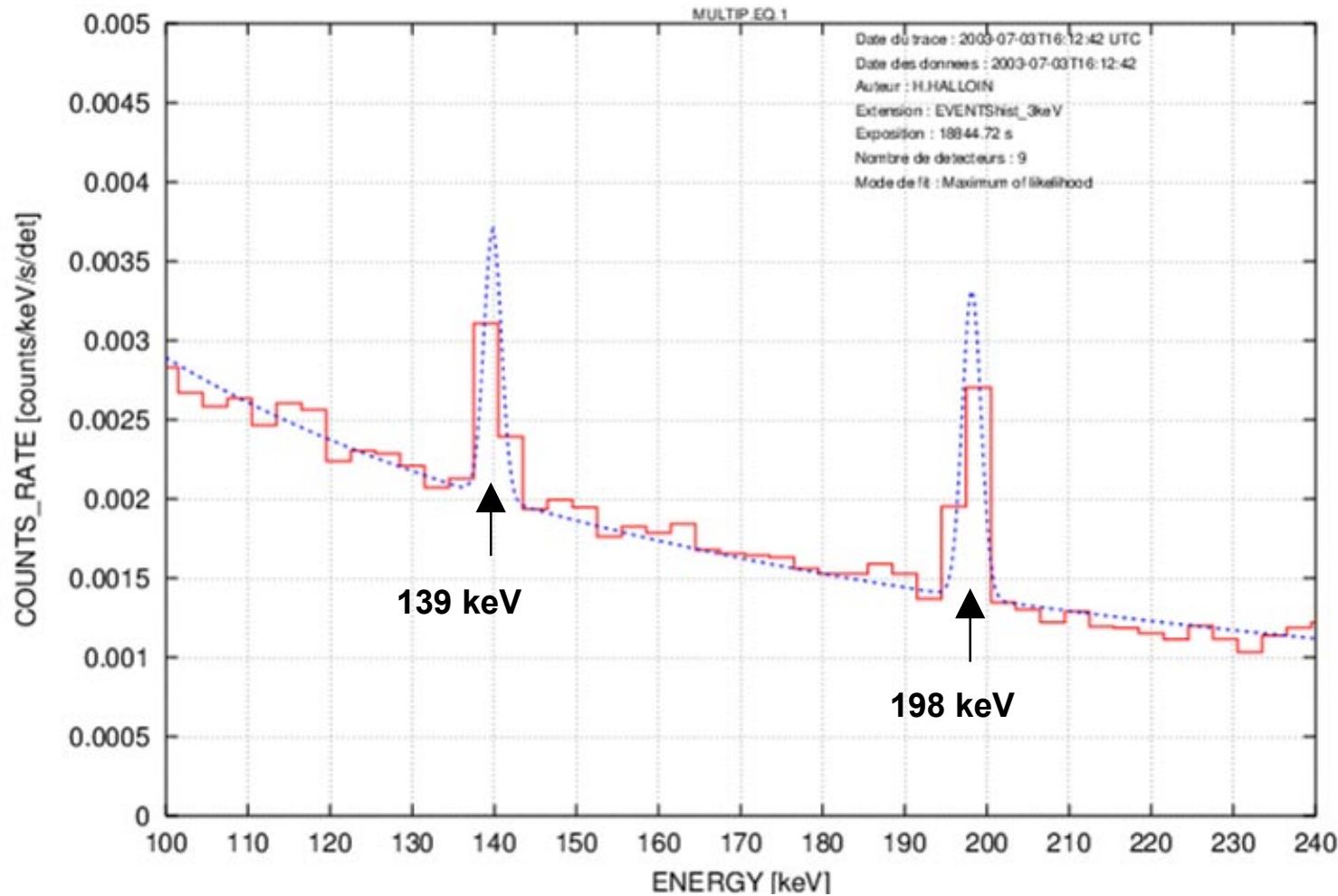
CLAIRE 2001 : carte du vol 14 juin 2001



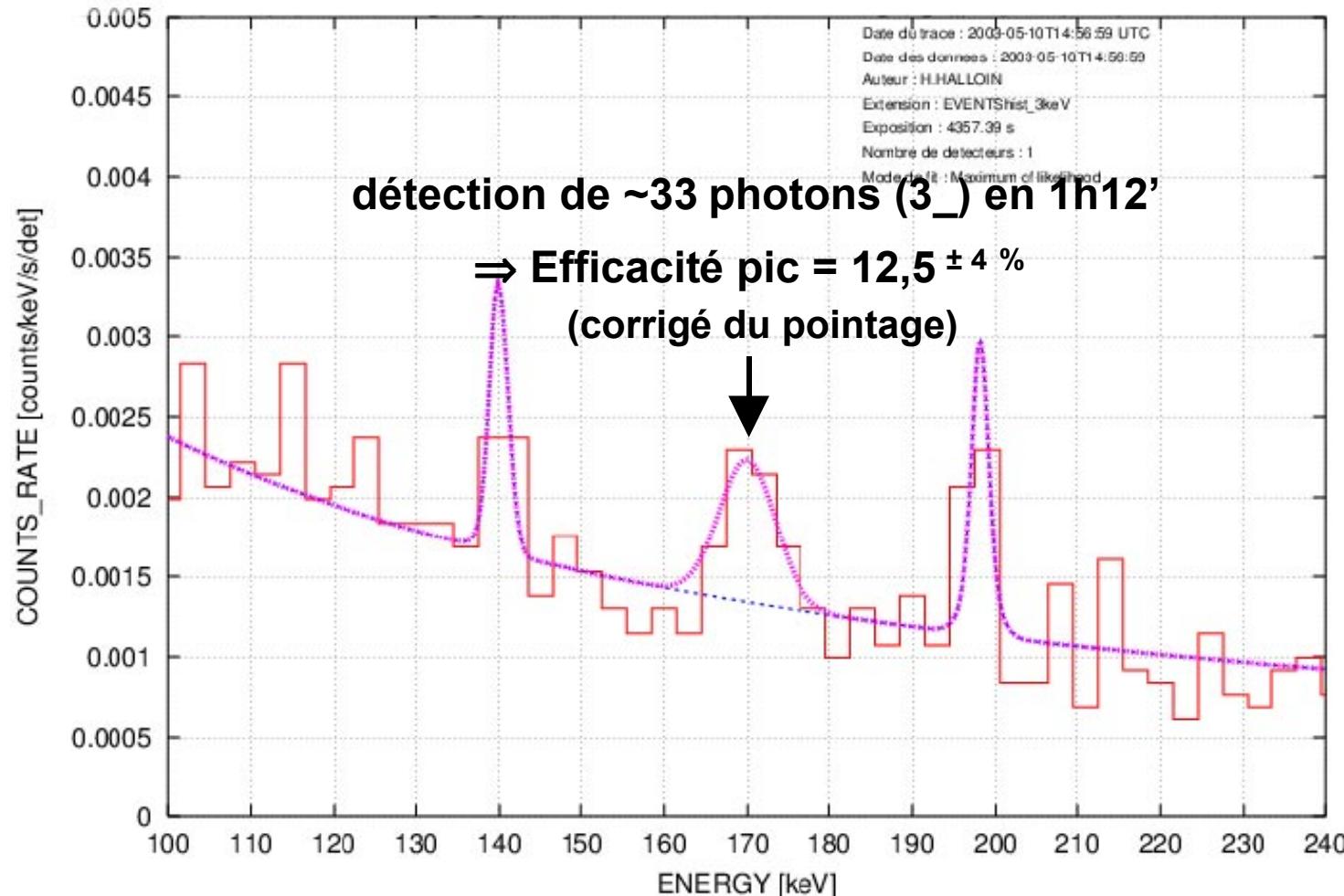
CLAIRE 2001 : pointages primaire et fin



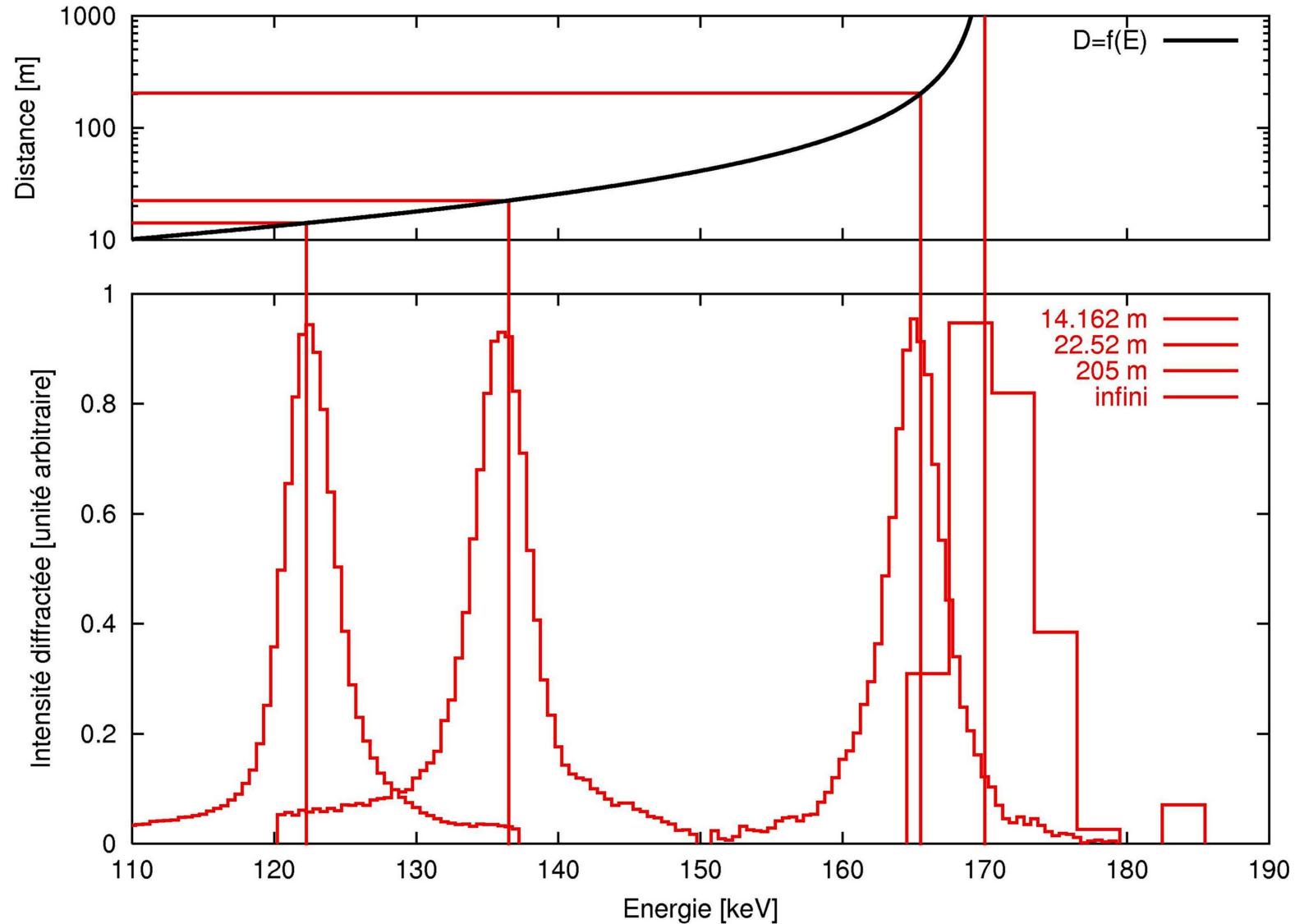
CLAIRE 2001 : bruit de fond au plafond (évts simples)



CLaire 2001 : première lumière d'une source astrophysique



CLAIRE 2001 : 14 m, 22.5 m, 205 m ... infini !



l'astronomie gamma
commence à voir

CLAIRE

