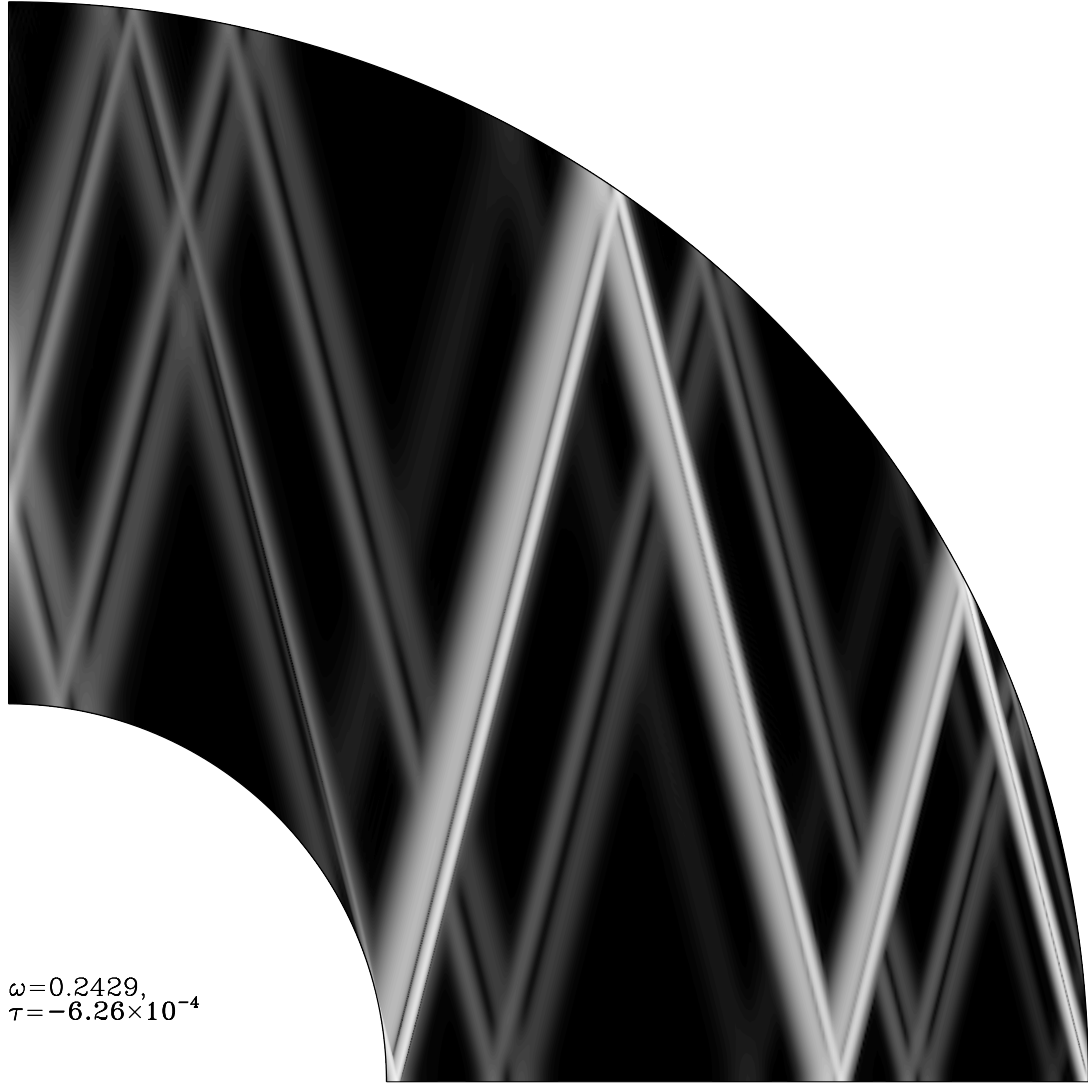


ERRATUM

In the paper “Ekman Layers and the Damping of Inertial r -Modes in a Spherical Shell: Application to Neutron Stars” by Michel Rieutord (ApJ, 550, 443 [2001]), Figure 2 was erroneously reproduced as Figure 1. The correct Figure 1 appears below. The Press sincerely regrets this error.



$$\omega=0.2429,$$

$$\tau=-6.26\times 10^{-4}$$

Mode 0^+ $\eta=0.35$ $L=1300$ $Nr=450$ $E=2.0\times 10^{-9}$ $CL=ff$

FIG. 1.—Kinetic energy distribution in a meridional plane of an inertial mode in a spherical shell associated with an equatorial attractor. A coexisting polar attractor is also slightly excited. The mode is axisymmetric with equatorial symmetry. Stress-free boundary conditions have been used on both shells; this solution was computed with an Ekman number of 2×10^{-9} and required 1300 spherical harmonics and 450 radial grid points (Gauss-Lobatto). The ratio of the inner radius to the outer radius is $\eta = 0.35$. $\omega = 0.2429$ and the damping rate $\tau = -6.26 \times 10^{-4}$ are given in dimensionless units as eq. (2). [See the electronic edition of the Journal for a color version of this figure.]